

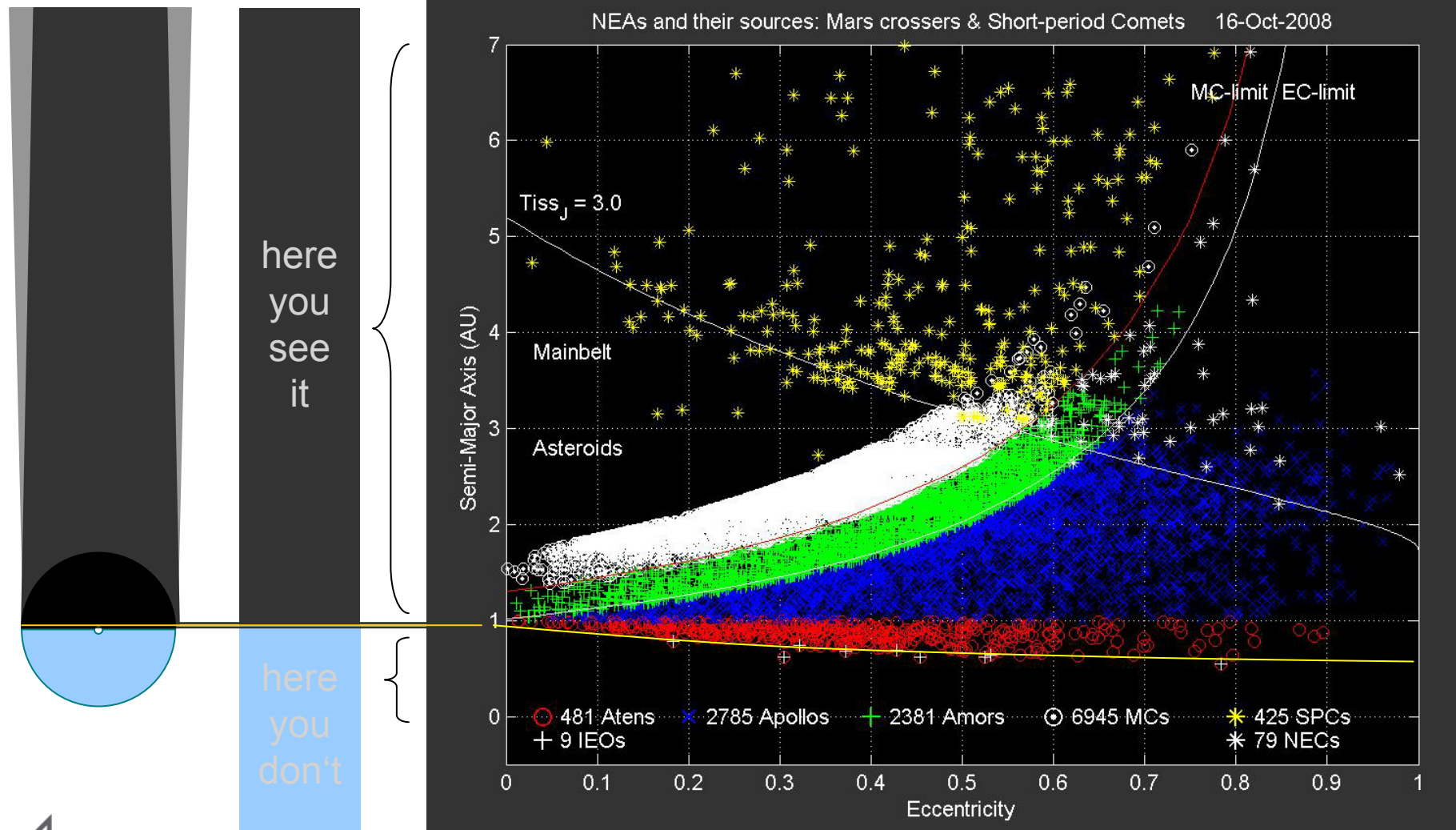


AsteroidFinder: Unveiling Objects Interior to Earth's Orbit

DLR German Aerospace Center
Bremen

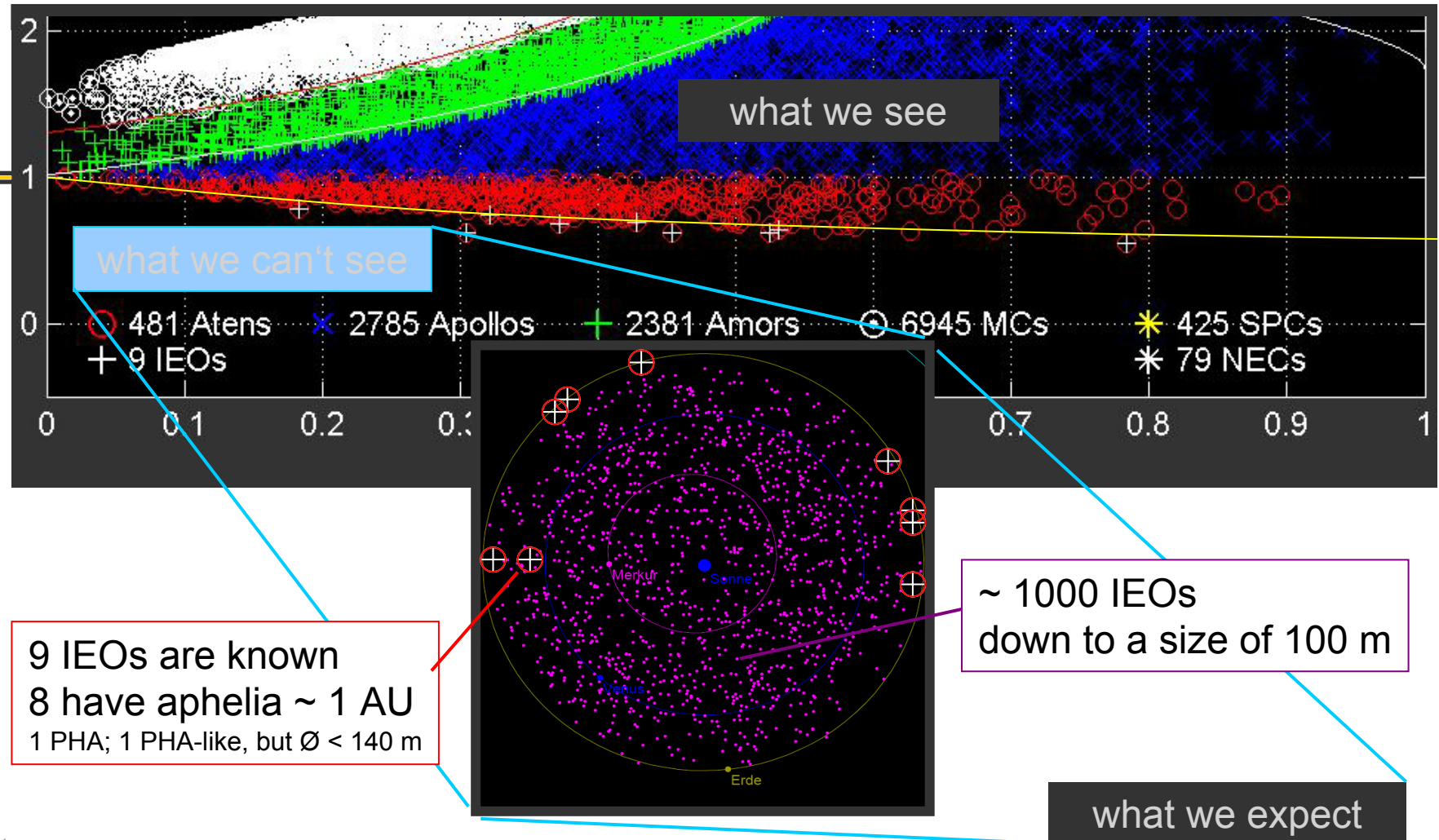


NEO Mitigation? - Shoot later. Ask first.

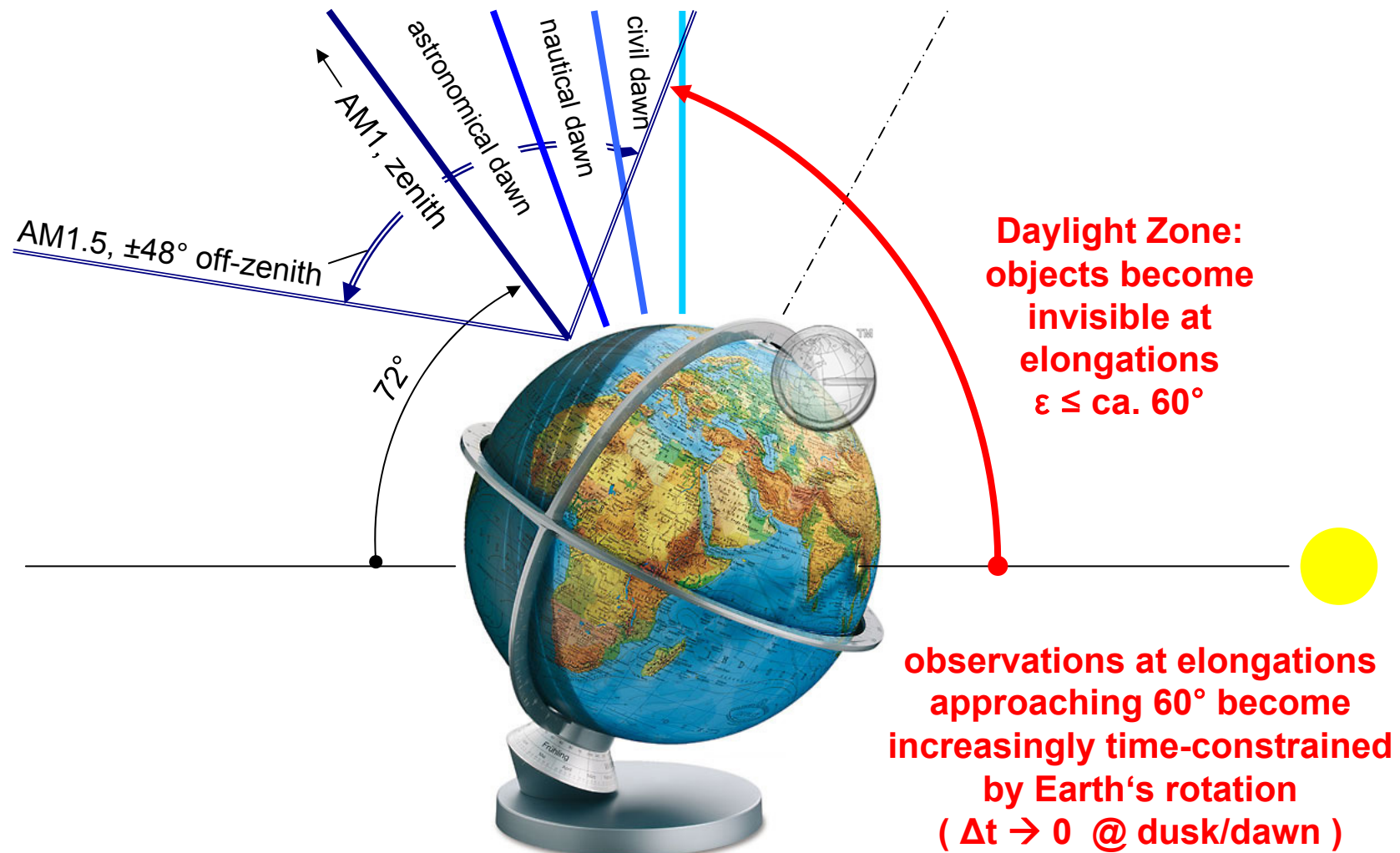




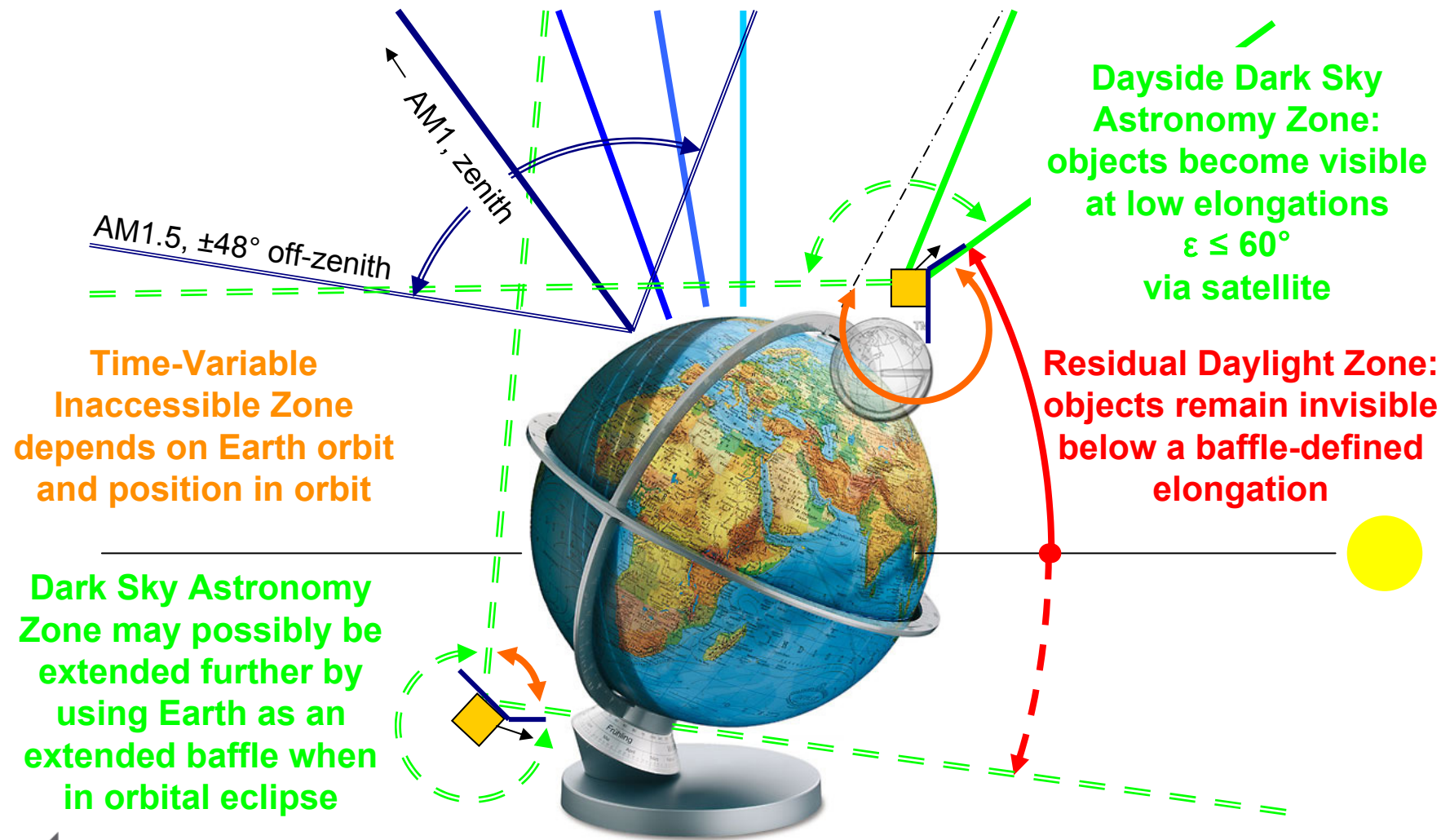
... and keep asking: Can I see them at all?!



the sky is the limit



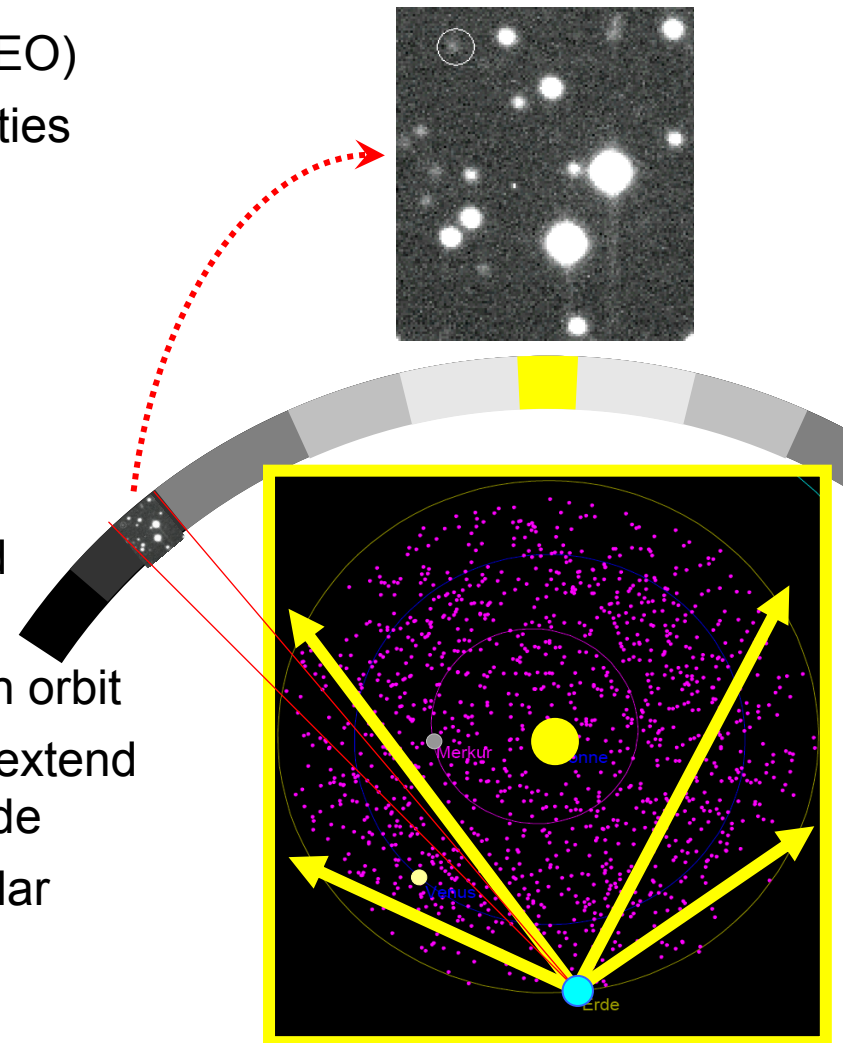
above the sky: AM0 on a sunny day





the task: unveiling the planets at day

- find objects Interior to Earth's Orbit (IEO)
- determine and catalogue their properties
 - population
 - orbital properties
 - size-frequency distribution
 - classes, groups, and families
- make use of the non-IEO background for night-like dayside astronomy
 - detect and track objects in Earth orbit
 - detect and track non-IEOs and extend known orbital arcs on the dayside
 - detect and monitor variable stellar objects (stars, supernovae,...)
 - monitor diffuse background

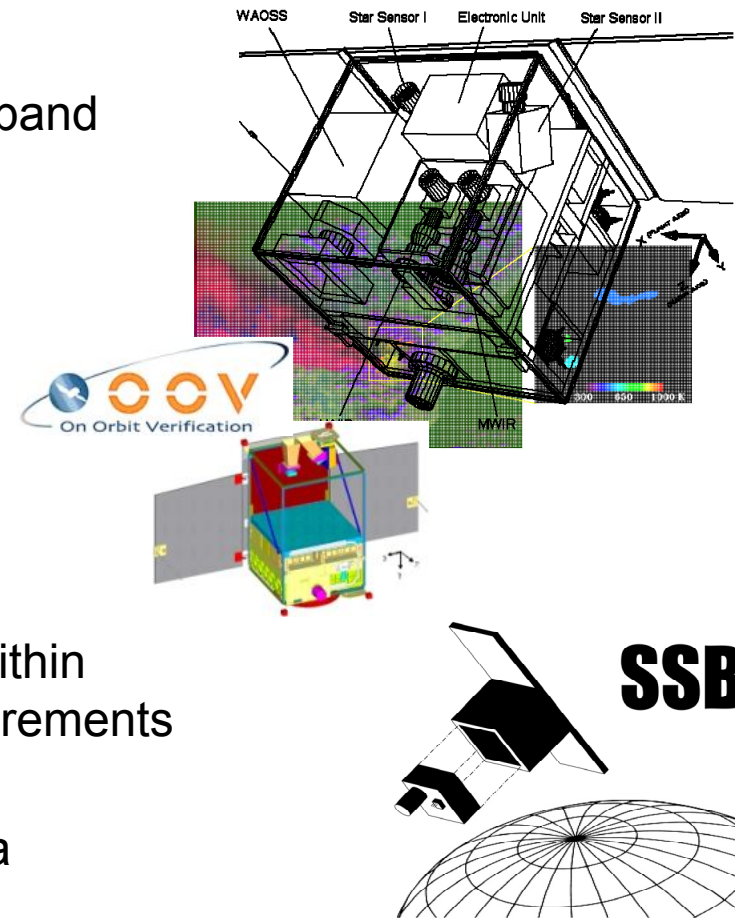




the foundation: 7 years of BIRD

the framework: „Kompaktsatellit“ programme

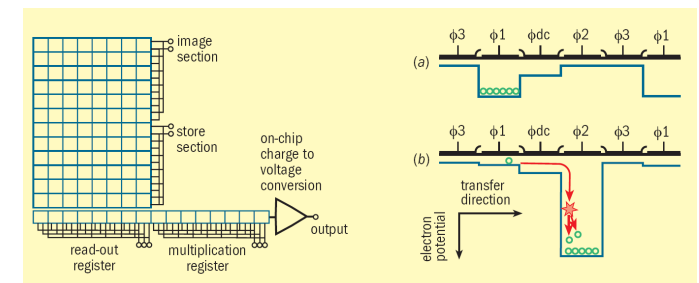
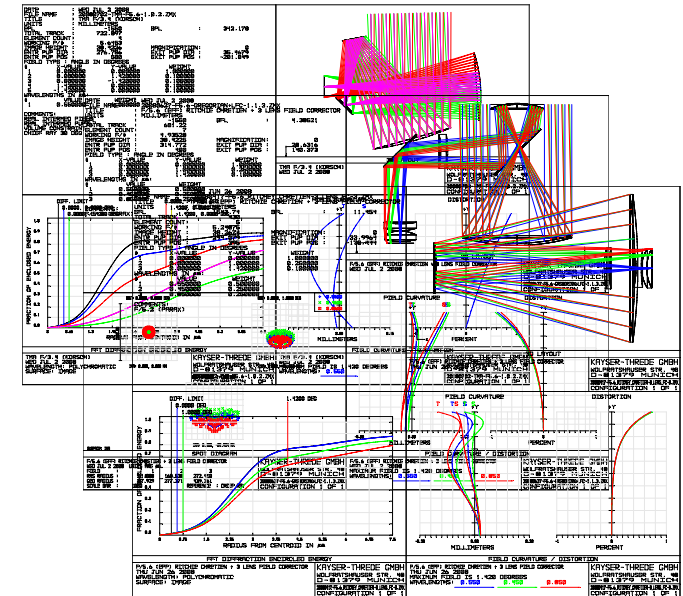
- *Bi-spectral InfraRed Detection*
 - optical bench for pixel-co-aligned dual band near, mid, thermal IR cameras
 - design life 1 year
 - reduced operability after 2 ¼ years
 - remains recoverable after 7 years
-
- Kompaktsatellit: best-of micro and mini satellites in a new 100...150 kg class
 - designed for scientific payloads from within DLR, on a bus tailored to science requirements
 - evolution towards a flight-proven set of subsystems and design variations, as a Standard Satellite Bus kit, /SSB



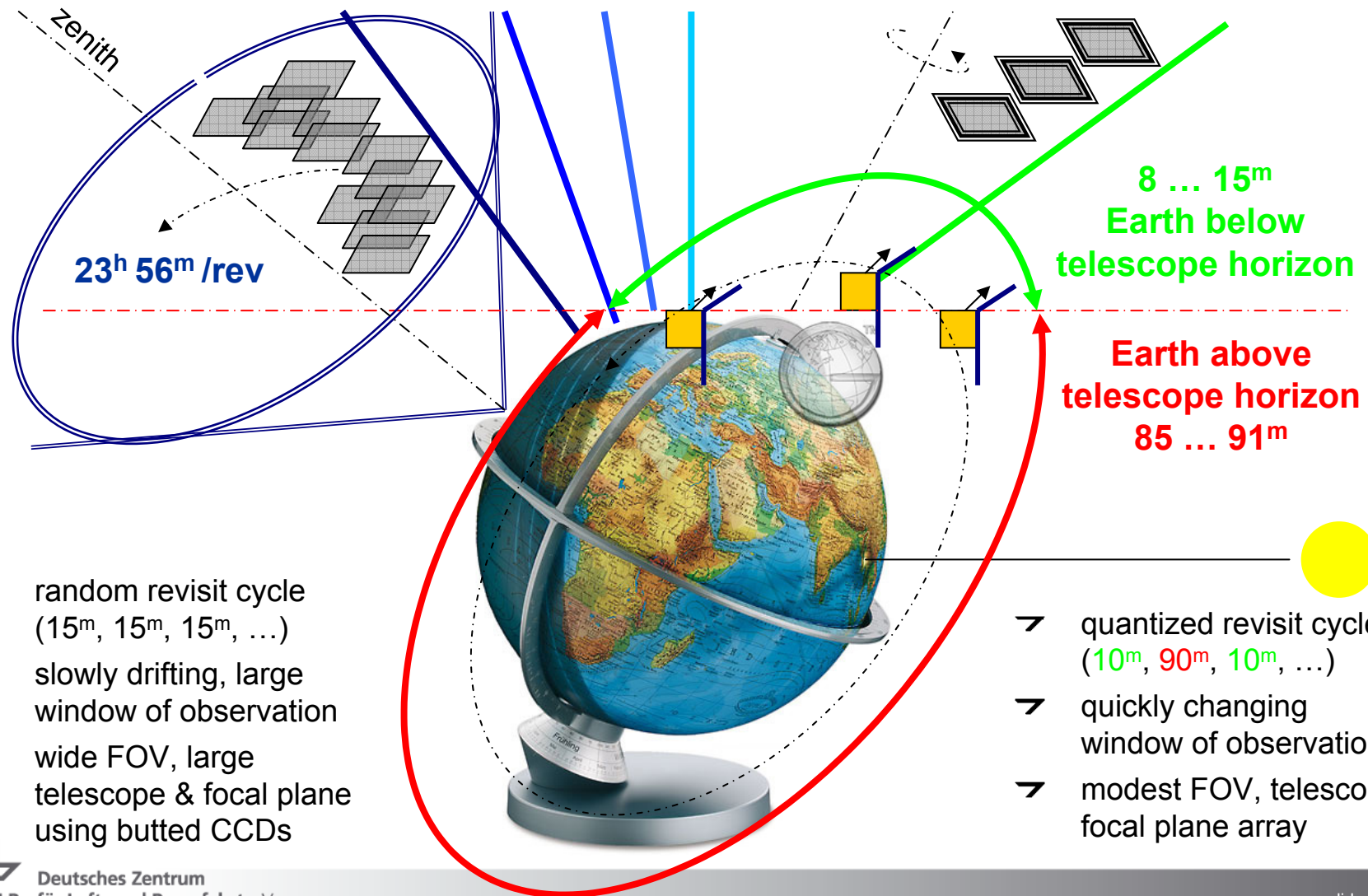


the instrument: advanced optics and electronics

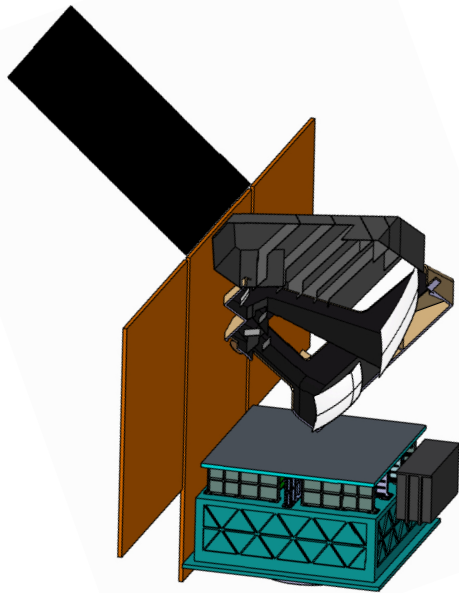
- limiting magnitude $V \geq 18.5 \text{ mag @ } 1 \text{ min}$
- FOV = $2^\circ \times 2^\circ$, astrometric accuracy $\sim 1 \dots 1.5''$
- off-axis telescope design
 - high straylight suppression
 - Sun : asteroid $\sim 10^{18} : 1$
 - planet : asteroid $\sim 10^8 : 1$
 - asteroid : background $\sim 4 : 1$
- electron-multiplied CCD sensors
 - fast read-out noise suppression
 - high flexibility in exposure timing and signal processing
 - registered stacking of $\sim 5 \text{ frames/s}$ to remove cosmics and satellite jitter
- area coverage & resolution \rightarrow high data rate



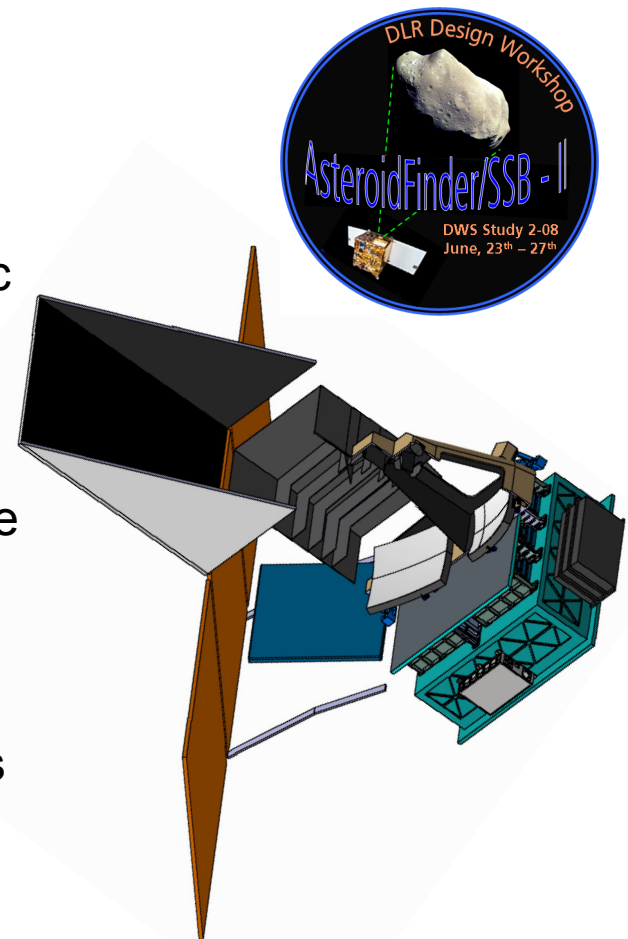
from the ground up: motion detection & orbital motion



taking shape: deep exposures in the light of day



- detecting sub-km asteroids at low elongations is feasible
- advanced design of the scientific payload combined with proven small satellite technology enables a cost-effective solution
- satellite orientation relative to the Sun and Earth is strongly coupled to thermal and optical sensitivities
- thorough end-to-end simulations are required very early in the project to iterate towards the best-performing overall design





Questions?



Who knows whether, when a comet
shall approach this globe to
destroy it, as it often has been
and will be destroyed, men will
not tear rocks from their
foundations by means of steam,
and hurl mountains, as the
giants are said to have done,
against the flaming mass? - And
then we shall have traditions of
Titans again, and of wars with
Heaven.

Lord Byron, 1822



Deutsches Zentrum
für Luft- und Raumfahrt e.V.
in der Helmholtz-Gemeinschaft

image credits: J.Hartmann M.Lieder DLR, G. Hahn DLR-EARN, ESA,NASA, astronautix.com, Sandia Labs, Kayser-Threde, geosmile.de; partially obtained via wikipedia

slide 11

AsteroidFinder/SSB > JB > RY-OR HB jtg > 16 Oct 2008



Supplementary Slides: Asteroid 101

- The Sky is the Limit – in Detail
- Space is not Unlimited
- Motion Detection
- Summed Up



Asteroid 101: The Sky is the Limit – in detail

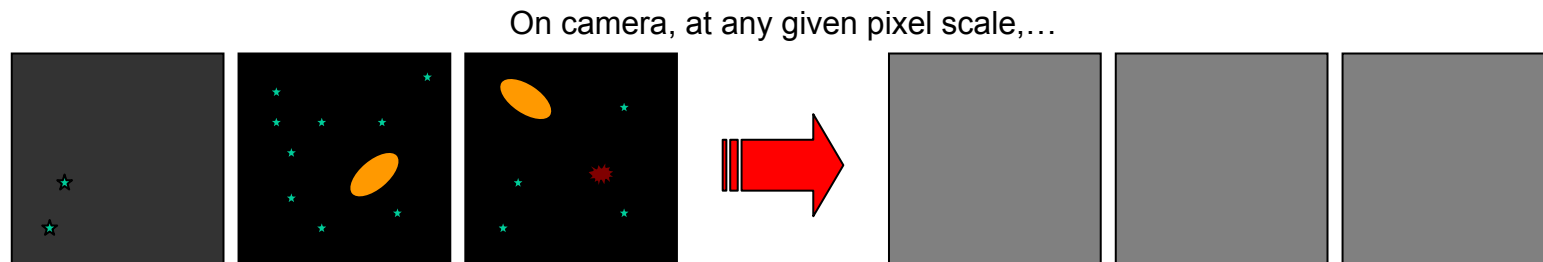
➤ Weather	clouds, haze, water vapour	< 18 km
➤ Blue Sky	scattered sunlight and moonlight	< 80 km
➤ Sky Glow	emission lines of excited molecules	< 250 km *
➤ SatelLight	gas discharge around spacecraft	< 500 km
➤ Air Glow	faint equatorial aurora	< 700 km *
➤ Aurora	bright polar aurora oval	< 1000 km *

➤ *) upper limit dependent on solar activity – **2011/12 is solar maximum!**



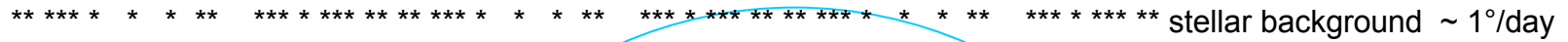
Asteroid 101: Space is not Unlimited

- stars and nebulae form a distant diffuse background at any resolution (“Billions and Billions”)
- interplanetary dust forms a local background that moves around the Sun (Zodiacal light, Lunar L4/5 dust clouds)
- the corona forms a variable background centered on the Sun, even beyond the area out to 32 solar radii covered by SOHO LASCO C-3



...diffuse background, stellar background, or passing asteroid may...

...*READ EXACTLY THE SAME*

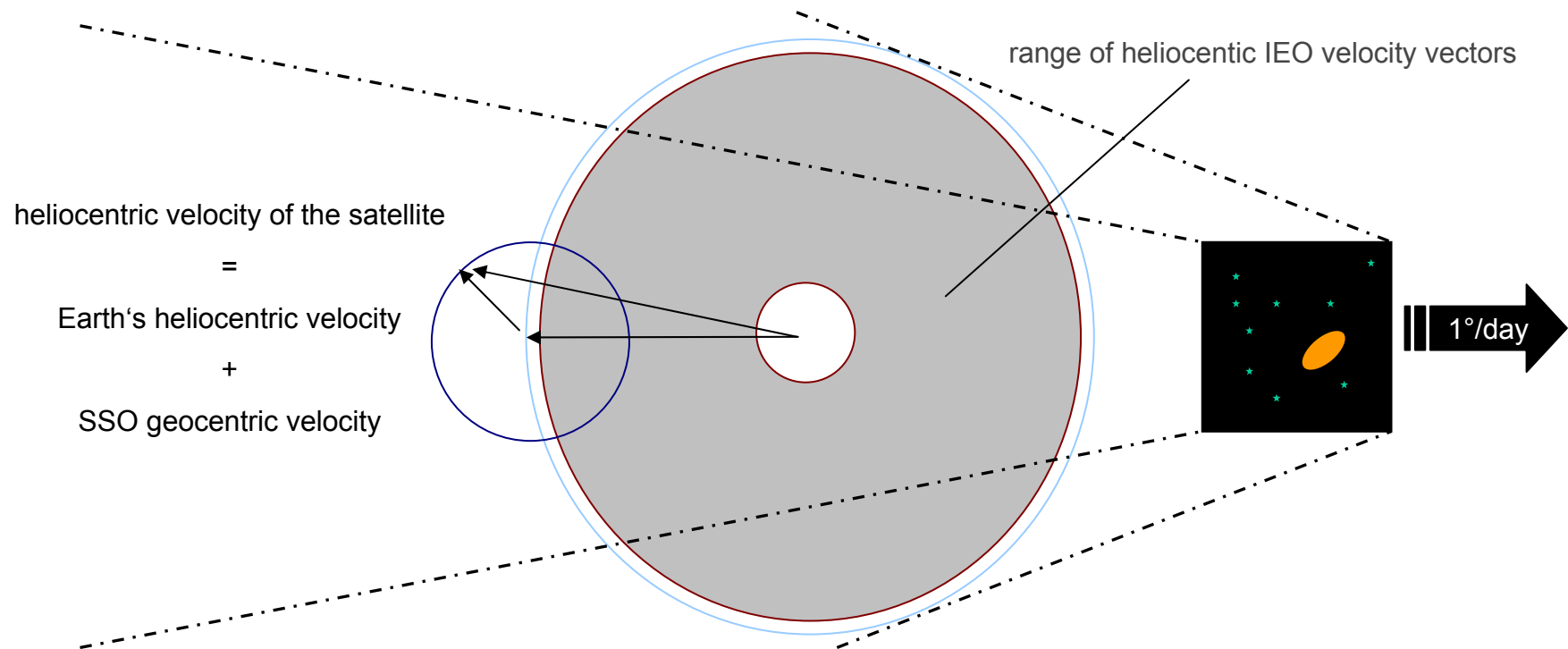


1: Δv

stellar background $\sim 1^\circ/\text{d}$

The diagram illustrates the orbital configuration for an Interplanetary Earth Orbiter (IEO). At the center is the Sun, represented by a yellow star. A large black circle represents Earth's orbit, with a radius of 0.983 AU. A smaller black ellipse represents the IEO's orbit, with a semi-major axis of 0.005 AU and an apocenter at 0.983 AU. The IEO is shown as a small blue satellite in its elliptical orbit. A cyan circle represents the Earth's orbit, with a radius of 0.983 AU. The Earth is shown as a blue sphere with a black satellite in its orbit. The Earth's orbit is labeled 'Earth in 0.983...1.017 AU elliptical orbit with satellite in 650...850 km SSO'. The IEO's orbit is labeled 'IEO in 0.005...0.983 AU elliptical orbit'. The diagram also shows the Earth's orbit as a cyan circle with a radius of 0.983 AU. The Earth is shown as a blue sphere with a black satellite in its orbit. The Earth's orbit is labeled 'Earth in 0.983...1.017 AU elliptical orbit with satellite in 650...850 km SSO'. The IEO's orbit is labeled 'IEO in 0.005...0.983 AU elliptical orbit'. The diagram also shows the Earth's orbit as a cyan circle with a radius of 0.983 AU. The Earth is shown as a blue sphere with a black satellite in its orbit. The Earth's orbit is labeled 'Earth in 0.983...1.017 AU elliptical orbit with satellite in 650...850 km SSO'. The IEO's orbit is labeled 'IEO in 0.005...0.983 AU elliptical orbit'.

Asteroid 101: Δv projected



- Zero relative velocities and angular rates are possible, with a few to a few hundred arcseconds/minute being typical
- Impossible to catch all at any time



Asteroid 101: easily over-optimized

- SNR: Keep area covered by pixel tiny
- Yield: Keep area covered by telescope huge
- Catch all sizes: Keep the shutter open for a long time
- Catch all orbits: Watch again and again

➤ *Get yourself a huge data volume*



Asteroid 101: ...Finder/SSB summed up

- detector: $(1K)^2$, ≤ 16 bit/pixel, ≥ 58 ms exposure, EMCCD, no filters
 - $(2^\circ)^2$ FOV, 4 independent channels, pixel scale $\sim 3''$ / pixel
 - motion detection threshold 2..5 pixels
 - $\sim 10''$ / >100 min \leq angular motion detection limit $\leq \sim 2^\circ.4$ / 1 min
 - expected yield: ~ 40 IEO's and ≥ 115000 Gbit raw data per year
 - recently, a dust ring just outside the orbit of Venus was discovered in Helios solar probe single pixel photomultiplier sensor raw data from 1974
- *Raw data are and remain valuable!*
- such as... main belt asteroids, Kuiper belt objects, supernovae, GRB afterglows, interplanetary dust background, exo-planet occultations, variable stars, fast-moving stars, and stellar background in general...



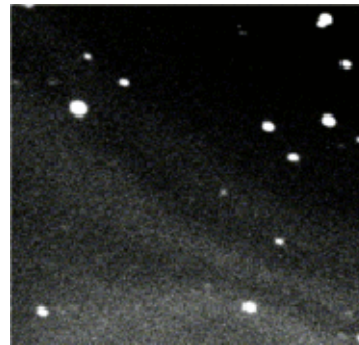
Asteroid 101: IEOs, NEOs, Mitigation

FAQ resources

- Gerhard Hahn, DLR EARN asteroid database: <http://earn.dlr.de/nea/> (*provides population graph in slides #2,3*)
- IAU: Minor Planet Center – Lists and Plots: Minor Planets: <http://cfa-www.harvard.edu/iau/lists/MPLists.html>
- NEODyS Near Earth Objects Dynamic Site: <http://newton.dm.unipi.it/cgi-bin/neodys/neoibo>
- Don Yeomans, NASA NEO Program – Current Impact Risks: <http://neo.jpl.nasa.gov/risk/>
- David Vokrouhlický, Paolo Farinella and William F. Bottke, Jr.; The Depletion of the Putative Vulcanoid Population via the Yarkovsky Effect, Icarus Volume 148, Issue 1, Nov. 2000, p. 147-152 (*google by title*)
- Patrick Michel, Vincenzo Zappalà, Alberto Cellino, Paolo Tanga; Estimated Abundance of Atens and Asteroids Evolving on Orbits between Earth and Sun, Icarus Volume 143, Issue 2, Feb. 2000, p. 421-424 (*google b.t.*)
- William F. Bottke, Jr., Alessandro Morbidelli, Robert Jedicke, Jean-Marc Petit, Harold F. Levison, Patrick Michel and Travis S. Metcalfe; Debaised Orbital and Absolute Magnitude Distribution of the Near-Earth Objects, Icarus Volume 156, Issue 2, Apr. 2002, p. 399-433 (*provided population data - google by title*)
- Tunguska Home Page, University of Bologna: <http://www-th.bo.infn.it/tunguska/> → Publications
- Michael J.S. Belton, Thomas H. Morgan, Nalin H. Samarasinha, Donald K. Yeomans (ed.), Mitigation of Hazardous Comets and Asteroids, Cambridge University Press, 2004
- Spaceguard Foundation: <http://spaceguard.rm.iasf.cnr.it/SGF/INDEX.html> <http://www.spaceguarduk.com/>
- Chrisian Gritzner, Kometen und Asteroiden – Bedrohung aus dem All, Aviatic Verlag (1999)
- Ralph Kahle, Modelle und Methoden zur Abwendung von Kollisionen von Asteroiden und Kometen mit der Erde, Doctoral Thesis, Technische Universität Berlin (2005):
http://opus.kobv.de/tuberlin/volltexte/2005/1127/pdf/kahle_ralph.pdf , this and more at
<http://www.weblab.dlr.de/rbrt/Publications/PubKahle.html>
- Jan Thimo Grundmann, Betrachtung des Missionsszenarios zur Verhinderung von Einschlägen von Asteroiden auf die Erde unter Berücksichtigung des Bedrohungspotentials und der technischen Möglichkeiten, diploma thesis, RWTH Aachen (2006): <http://www.kiwikommando.de/space4space/> (*provisional*)



Asteroid 101: The Devil is in the Details



(99942) Apophis

...named after the Ancient Egyptian Uncreator who dwells in the eternal darkness of the underworld. A close Earth flyby on Fri 13 Apr 2029 below geostationary altitude will gravity-assist Apophis for anything between a ~0.1 AU miss and a dead centre Earth impact on 13 Apr 2036, at 2.2×10^{-5} estimated probability.

Hint: (99942) + ✂ + ⌕ → '666' + 42 :-)